Theoretical Studies on the Complex Organic Molecules Origin of the Mid Infrared UIE bands in Planetary Nebulae

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## **UIE Bands in Young PN**

Understanding the Deep Space Chemistry



Despite their harsh environments, PNs can be the host of complex organic molecules Sun Kwok, Astrophys Space Sci, 319:5 (2009)

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## What Do UIE Bands and Plateaus Tell Us

#### Origin UIE bands are broad

• Collection of atoms in forms of clusters or molecules

#### Size

#### Plateau features

 UIEs sit on top of the very broad plateau features. By assuming the same source for both UIEs and plateaus, we can estimate the size of more than 100 atoms for UIE carrier

#### Bonds

#### UIE bands are discrete

 If the UIE carrier is one big molecules or group of molecules, they should composed of just a few types of bonds. Usually the wavelength range of 3-10 µm is the bond type control region, wavelength longer than 10 µm is stereochemistry control region.



# An Example of Broad IR Features

FT-IR Spectrum of a Liquid Crystal Polymer Known as FLCP



#### **UIE Carriers as Seen by Mathematics**

#### **Spectral Similarity**

 $X = \{x_1, x_2, ..., x_i, ..., x_n\}$  target molecule intensity or wavelengths  $Y = \{y_1, y_2, ..., y_i, ..., y_n\}$  reference molecule intensity or wavelengths

Properties = f(strcutue)





$$i_{p} = \frac{\sum_{i} (x_{i} - \overline{x}) \cdot (y_{i} - \overline{y})}{\sqrt{\sum_{i} (x_{i} - \overline{x})^{2} \cdot \sum_{i} (y_{i} - \overline{y})^{2}}}$$

Pearson's product moment

correlation coefficient.

 $r_p=1$  prefect linear correlation or similarity.

For performing similarity we need
1- a pool of reference spectra with same
wavelength resolution and normalized intensities
with respect to the strongest peak.
2- the unknown spectra should be prepared to match
The conditions of reference spectra.



#### UIE Carriers as Seen by Mathematics Spectral Similarity

Theoretical database with more than 275,000 molecules up to the mass of 500 a.m.u Calculated IR spectra at EDF2/6-31G\* and ωB97X-D/6-31G\* EDF2 stands for Empirical density functionals

 $E^{EDF2} = a_1 E^{HF} + a_2 E^{LSDA} + a_3 E^{BLYP} + a_4 E^{B3LYP} + a_5 E^{EDF1}$ 

 $a_i$ : is determined by least-square fit of HF, LSDA, BLYP, B3LYP and EDF1 harmonic frequencies to experimental data

 $E_{XC}^{EDF2} = b_1 E_X^{F30} + b_2 E_X^{D30} + b_3 E_X^{B88} + b_4 E_X^{EDF1} + b_5 E_C^{VWN} + b_6 E_C^{LYP} + b_7 E_C^{EDF1}$ 

Overall RMS:  $34 \ cm^{-1}$ ,

Target: Experimental IR gas phase (NIST) Reference: EDF/6-31G\* data base

Total: 35 tries (known Spectra) Successful exact hit by Similarity: 13 Reliability : 37%

Lin, C. Y.; George, M. W.; Gill, P. M. W., *Australian Journal of Chemistry* 2004, *57* (4), 365-370. Adamson, R. D.; Gill, P. M. W.; Pople, J. A., *Chemical Physics Letters* 1998, *284* (1–2), 6-11.



### **UIE Carriers as Seen by Mathematics**

Spectral Similarity results for IRAS2128+5050



C<sub>15</sub>H<sub>13</sub>N<sub>3</sub>O 4-mthyl-n-(1h-pyrrolo[2,3-b]pyridine-1-yl]benzenecarboxamide



#### **Quantum Mechanics & Quantum Chemistry**



Guzman and Bader, Coord.Chem.Review, 249:633-662 (2005)



## **Applications of Quantum Chemistry**

Direct modeling of the UIE bands carriers.

Understanding the vibrational origin of the UIE bands.

Search for the new species



## **Direct Modeling of the UIE Bands**

#### Adding Aliphatic Groups to the PAH Core



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## **Direct Modeling of the UIE Bands**

Mixed Aliphatic-Aromatic Organic Nanoparticles (MAONs)



We try to understand which fragments inside such a big molecule are responsible for the observed IR features. Then we can return back to the observations and track the existences of such fragments.

Sadjadi, Kwok and Zhang, Journal of Physics, Conference Series 728 (2016) 062003

## **Direct Modeling of the UIE Bands**

**Amorphous Hydrocarbon** 



#### Wavelength (µm)

Amorphous hydrocarbon shows a very complex bonding pattern, this makes the interpretation of IR spectrum very difficult. Here the advanced bonding model via quantum theory of atoms in molecules (QTAIM) is applied to reveal the nature of bonds.



### The Origin of the UIE Bands

3.28 and 3.3 µm components of 3 µm feature



986 normal modes



Wavelength (µm)



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### The Origin of the UIE Bands

Vibrational analysis on 1248 modes within the range of 3.3-3.4  $\mu$ m





### The Origin of the UIE Bands

Vibrational analysis within the range of 11-15 µm



Sadjadi, Zhang and Kwok, ApJ, 807:95 (2015)



#### **Search for New Species**

**Molecules with Olefinic Functional Group** 



We provided an alternative new explanation that 6.0 µm feature is not due to CO but due to olefinic containing compounds (-C=C-) Hsia, Sadjadi, Zhang and Kwok, ApJ,832:213 (2016)



#### **Search for New Species**

**Fulleranes** 



 $C_{60}H_{18}$  as one of the most kinetically and thermodynamically stable fulleranes shows unique IR feature at 15 µm. We use this feature to search for this species in space. Zhang, Sadjadi, Hsia and Kwok, ApJ,845:76 (2017)

### **Our List of Questions for Subaru**

- What elements are these UIE carries made of?
- How many atoms they have?
- What are their molecular structures?
- What electric charges they carry (+, ++, ...)?
- What are their electronic states (Singlet, Doublet,...)?
- At what temperature and pressure are they formed?
- How do they form?
- Are they entirely formed by the gas phase reactions or condensed phase reactions?
- Where do they form?. At the center of the stars or other places?
- How can we get the confirmation/rejection of our understandings of these complex phenomena obtained in the mid-infrared region, in another region of spectra like optical or near infrared regions?



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